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# State reconstruction of wave packets moving in time-dependent potentials and the existence of Wronskian pairs 

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#### Abstract

We determine a class of time-dependent potentials that support the state reconstruction of one-dimensional wave packets. For this we extend the projection method (Leonhardt $U$ and Raymer M G 1996 Phys. Rev. Lett. 76 1985) to explicitly time-dependent situations. We require the existence of Wronskian pairs (regular and irregular wavefunctions) and analyse the consequences on the possible potentials.


Quantum-mechanical wave packets travelling in known but otherwise arbitrary potentials possess an interesting property: the motion of the packet reveals the quantum state [1-4]. The spatio-temporal probability distribution $\operatorname{pr}(x, t)$ of the corresponding particles contains sufficient information to retrieve the density matrix $\rho_{m n}$ of the wave packet (denoted in the energy representation). This theoretical concept [1-4] unifies and generalizes several experimental methods of state determination: Optical homodyne tomography [5-7], molecular fluorescence tomography [8, 9], and atomic-beam tomography [10]. Up to now, however, the concept [1-4] has been restricted to stationary potentials $\ddagger$. State reconstruction of wave packets travelling in time-dependent potentials is interesting, because the motion could increase in complexity (most notably in quantum chaos). Furthermore, the phase retrieval of nonlinear waves [12] is based on an effective time-dependent potential. This problem is particularly relevant [12] for the state determination of travelling Bose-Einstein condensates observed in phase-contrast imaging [13]. Can we extend the ideas developed for stationary potentials [1-4] to the time-dependent case [14]?

What are these ideas? The density matrix $\rho_{m n}$ in the energy representation is reconstructed as the average

$$
\begin{equation*}
\rho_{m n}=\left\langle\left\langle\frac{\partial}{\partial x}\left[\psi_{m}^{*}(x, t) \varphi_{n}(x, t)\right]\right\rangle\right\rangle_{x, t} \tag{1}
\end{equation*}
$$

with respect to the measured positions $x$ at all times $t$. By $\langle\langle F(x, t)\rangle\rangle_{x, t}$ we denote the average

$$
\begin{equation*}
\langle\langle F(x, t)\rangle\rangle_{x, t} \equiv \lim _{T \rightarrow \infty} T^{-1} \int_{-T / 2}^{+T / 2} \int_{-\infty}^{+\infty} \operatorname{pr}(x, t) F(x, t) \mathrm{d} x \mathrm{~d} t \tag{2}
\end{equation*}
$$

[^0]of a function $F(x, t)$ with respect to the spatio-temporal probability distribution $\operatorname{pr}(x, t)$. Here
\[

$$
\begin{equation*}
\psi_{n}(x, t)=u_{n}(x) \mathrm{e}^{-1 \omega_{n} t} \quad \varphi_{n}(x, t)=v_{n}(x) \mathrm{e}^{-1 \omega_{n} t} \tag{3}
\end{equation*}
$$

\]

denote the regular and irregular wavefunctions of the energy eigenstates $|n\rangle$. (For simplicity we assume a discrete spectrum. For the continuous case see [15].) Both the regular and the irregular wavefunctions are solutions of the Schrödinger equation

$$
\begin{equation*}
\mathrm{i} \dot{\phi}_{n}=-\frac{1}{2} \phi_{n}^{\prime \prime}+U \phi_{n} \tag{4}
\end{equation*}
$$

in appropriate units. Regular wavefunctions are normalized to unity, whereas irregular ones obey the Wronskian condition

$$
\begin{equation*}
W_{n}=u_{n} v_{n}^{\prime}-u_{n}^{\prime} v_{n}=2 \tag{5}
\end{equation*}
$$

The condition implies [3] that irregular wavefunctions must not be normalizable. Analytic examples of irregular wavefunctions are given in [16-18]. The key mathematical property of the spatial derivative $\left(\psi_{m}^{*} \varphi_{n}\right)^{\prime}$ of regular and irregular wavefunction is the fact [1-4] that these objects form an orthonormal system on products of wavefunctions $\psi_{\mu} \psi_{v}^{*}$, i.e.

$$
\begin{equation*}
D_{\mu \nu}^{m n} \equiv \lim _{T \rightarrow \infty} T^{-1} \int_{-T / 2}^{+T / 2} \int_{-\infty}^{+\infty} \psi_{\mu} \psi_{v}^{*}\left(\psi_{m}^{*} \varphi_{n}\right)^{\prime} \mathrm{d} x \mathrm{~d} t=\delta_{\mu m} \delta_{\nu n} \tag{6}
\end{equation*}
$$

Therefore, $\left(\psi_{m}^{*} \varphi_{n}\right)^{\prime}$ projects the density-matrix elements $\rho_{m n}$ out of the spatio-temporal probability distribution

$$
\begin{equation*}
\operatorname{pr}(x, t)=\langle x, t| \hat{\rho}|x, t\rangle=\sum_{\mu \nu} \rho_{\mu \nu} \psi_{\mu}(x, t) \psi_{\nu}^{*}(x, t) . \tag{7}
\end{equation*}
$$

State reconstruction of wave packets from $\operatorname{pr}(x, t)$ is a non-trivial property, because some counterexamples are known [19] for multi-dimensional spaces where the mapping between $\operatorname{pr}(x, t)$ and the quantum state is not unique. Therefore, to extend state reconstruction to the case of time-dependent potentials, we seek a similar structure as in the stationary case. There are no stationary states in general, but we might assume that a basis of states $|n\rangle$ exists that have pairs of regular and irregular wavefunctions $\psi_{n}(x, t)$ and $\varphi_{n}(x, t)$. What do we mean by that? Let us assume that the functions $\psi_{n}$ and $\varphi_{n}$ are solutions of the Schrödinger equation (4) that form a Wronskian pair

$$
\begin{equation*}
\psi_{n}^{*} \varphi_{n}^{\prime}-\psi_{n}^{\prime} \varphi_{n}^{*}=W_{n} \quad W_{n}^{\prime}=0 \tag{8}
\end{equation*}
$$

How to find the wavefunctions $\psi_{n}$ and $\varphi_{n}$ will not concern us at the moment; we will come to this point later. Let us first examine the implications of the sheer existence of Wronskian pairs. Given a function $\psi_{n}$, equation (8) formulates a differential equation for $\varphi_{n}$ with the solution

$$
\begin{equation*}
\varphi_{n}=W_{n} \psi_{n} \int \frac{\mathrm{~d} x}{\psi_{n}^{*} \psi_{n}} \tag{9}
\end{equation*}
$$

as is easily verified. The integration in the solution (9) is an undetermined integral with the set of integration constants corresponding to the set of unique solutions of the differential equation (8). Therefore, equation (9) is the general solution of (8). Equation (9) thus describes uniquely the relation between the two functions $\psi_{n}$ and $\varphi_{n}$ that form a Wronskian pair (8).

To proceed, we expand the wavefunctions $\psi_{n}$ in (real) amplitude and (real) phase functions

$$
\begin{equation*}
\psi_{n}=u_{n} \mathrm{e}^{\mathrm{i} S_{n}} \tag{10}
\end{equation*}
$$

We see from (9) that $\varphi_{n}$ has the same phase factor as $\psi_{n}$ :

$$
\begin{equation*}
\varphi_{n}=v_{n} \mathrm{e}^{\mathrm{i} S_{n}} \tag{11}
\end{equation*}
$$

and the amplitude

$$
\begin{equation*}
v_{n}=W_{n} u_{n} \int \frac{\mathrm{~d} x}{u_{n}^{2}} \tag{12}
\end{equation*}
$$

Let us formulate the Schrödinger equation (4) in terms of amplitude and phase:

$$
\begin{align*}
& -\frac{1}{2} u_{n}^{\prime \prime}+\left(U+\frac{1}{2} S_{n}^{\prime 2}+\dot{S}_{n}\right) u_{n}=0  \tag{13}\\
& \dot{u}_{n}+\frac{1}{2} S_{n}^{\prime \prime} u_{n}+S_{n}^{\prime} u_{n}^{\prime}=0 \tag{14}
\end{align*}
$$

(and identical equations for $v_{n}$ ). Equation (13) plays the role of the stationary Schrödinger equation in the explicitly time-dependent case. We see that the potential $U$ is modified by the term $\frac{1}{2} S_{n}^{\prime 2}+\dot{S}_{n}$ that contains the spatial structure and the evolution of the phase $S_{n}$. Equation (14) is a form of the conservation law for the probability density $u_{n}^{2}$ with the probability flux $S_{n}^{\prime} u_{n}^{2}$, because from (14) we obtain the result that $\partial\left(u_{n}^{2}\right) / \partial t+\left(S_{n}^{\prime} u_{n}^{2}\right)^{\prime}$ vanishes.

We show in appendix A that $v_{n}$ satisfies the Schrödinger equation (13) automatically, whereas the probability conservation (14) seriously restricts the possible choice of the phase functions $S_{n}$

$$
\begin{equation*}
S_{n}^{\prime \prime}=-\frac{1}{2} \frac{\dot{W}_{n}}{W_{n}} \tag{15}
\end{equation*}
$$

The phase $S_{n}$ is a quadratic function in space

$$
\begin{equation*}
S_{n}=a_{n}(t) \frac{x^{2}}{2}+b_{n}(t) x-\Omega_{n}(t) \tag{16}
\end{equation*}
$$

Given the structure (16) of the phase $S_{n}$, we utilize the probability conservation (14) to determine the structure of the amplitudes $u_{n}$. We show in appendix B that $u_{n}(x, t)$ is a scaled function $w_{n}(\xi)$ such that

$$
\begin{array}{ll}
u_{n}=\eta^{-1 / 2} w_{n}(x / \eta-\zeta) \\
a_{n}=\eta^{-1} \dot{\eta} \quad b_{n}=\eta \dot{\zeta} \tag{18}
\end{array}
$$

The irregular wavefunctions obey the same scaling.
We have seen that the existence of Wronskian pairs does restrict the wavefunctions. Let us require another property that originates from the normalization

$$
\begin{equation*}
D_{m n}^{m n}=1 \tag{19}
\end{equation*}
$$

of the desired orthogonal system (6) on products of wavefunctions. From (10) and (11) we obtain
$D_{m n}^{m n}=\lim _{T \rightarrow \infty} T^{-1} \int_{-T / 2}^{+T / 2} \int_{-\infty}^{+\infty} u_{m} u_{n}\left[\left(u_{m} v_{n}\right)^{\prime}+\mathrm{i}\left(S_{m}^{\prime}-S_{n}^{\prime}\right) u_{m} v_{n}\right] \mathrm{d} x \mathrm{~d} t$.
In order to guarantee that $D_{m n}^{m n}$ is real we require that $\left(S_{m}^{\prime}-S_{n}^{\prime}\right)$ be zero, i.e.

$$
\begin{equation*}
S_{n}=S(x, t)-\Omega_{n}(t) \tag{21}
\end{equation*}
$$

From (16) and (18) we obtain

$$
\begin{equation*}
S(x, t)=\frac{\dot{\eta}}{\eta} \frac{x^{2}}{2}+\eta \dot{\zeta} x \tag{22}
\end{equation*}
$$

This result restricts the class of potentials $U(x, t)$. To see this we define the function

$$
\begin{equation*}
F_{n} \equiv U+\frac{1}{2} S^{\prime 2}+\dot{S}-\dot{\Omega}_{n} \tag{23}
\end{equation*}
$$

and show in appendix C that $F_{n}$ obeys the scaling

$$
\begin{equation*}
F_{n}=\eta^{-2} G_{n}(x / \eta-\zeta) \tag{24}
\end{equation*}
$$

Therefore we are lead to the requirement that the spatially independent part of $F_{n}$ scales as

$$
\begin{equation*}
\dot{\Omega}_{n}(t)=\frac{\omega_{n}}{\eta^{2}(t)} . \tag{25}
\end{equation*}
$$

The set of constants $\omega_{n}$ play the role of energy eigenvalues and will be determined soon. We obtain for the phases

$$
\begin{equation*}
S_{n}(x, t)=\frac{\dot{\eta}}{\eta} \frac{x^{2}}{2}+\eta \dot{\zeta} x-\omega_{n} \int \frac{\mathrm{~d} t}{\eta^{2}} . \tag{26}
\end{equation*}
$$

Furthermore, $F_{n}+\dot{\Omega}_{n}$ should scale as

$$
\begin{equation*}
F_{n}+\dot{\Omega}_{n}=U+\frac{1}{2} S^{\prime 2}+\dot{S}=\eta^{-2} V(x / \eta-\zeta) \tag{27}
\end{equation*}
$$

We use this scaling property and equation (22) to obtain the class of potentials that support a Wronskian pair (8) with the property (21):

$$
\begin{equation*}
U=\eta^{-2} V(x / \eta-\zeta)-\frac{\ddot{\eta}}{\eta} \frac{x^{2}}{2}-(2 \dot{\eta} \dot{\zeta}+\eta \ddot{\zeta}) x-\frac{\eta^{2} \dot{\zeta}^{2}}{2} \tag{28}
\end{equation*}
$$

This is the central result of our paper. The existence of Wronskian pairs (8) with the property (21) restricts the class of time-dependent potentials $U(x, t)$. Apart from a quadratic term they are scaled and shifted stationary potentials $V(\xi)$. The explicit time dependence is brought about by the scaling function $\eta(t)$ and the potentially time-dependent shift $\zeta(t)$.

We also employ the scaling property (27) to find the set of wavefunctions and the numbers $\omega_{n}$. The amplitudes $u_{n}$ obey (13) and, consequently,

$$
\begin{equation*}
-\frac{1}{2} u_{n}^{\prime \prime}+\eta^{-2} V(x / \eta-\zeta) u_{n}=\eta^{-2} \omega_{n} u_{n} \tag{29}
\end{equation*}
$$

We apply the scaling (17) of the wavefunctions and see that the $w_{n}(\xi)$ are the eigenfunctions of the stationary Schrödinger equation

$$
\begin{equation*}
-\frac{1}{2} \frac{\mathrm{~d}^{2} w_{n}}{\mathrm{~d} \xi^{2}}+V(\xi) w_{n}=\omega_{n} w_{n} \tag{30}
\end{equation*}
$$

with potential $V(\xi)$ and eigenvalues $\omega_{n}$. The irregular wavefunction are then given by (12).
So far, we were concerned about the implications of the existence of Wronskian pairs (8) in the case of a time-dependent potential. We determined the class of potentials that support such pairs. Finally, we prove that our mathematical objects are indeed useful for state reconstruction. We consider the overlap (6), utilize the structure (10) and (11) of the wavefunctions and the phases (26) to obtain
$D_{\mu \nu}^{m n}=\lim _{T \rightarrow \infty} T^{-1} \int_{-T / 2}^{+T / 2} \int_{-\infty}^{+\infty} u_{\mu} u_{v}\left(u_{m} v_{n}\right)^{\prime} \exp \left[\mathrm{i}\left(\omega_{m}-\omega_{n}-\omega_{\mu}+\omega_{\nu}\right) \theta\right] \mathrm{d} x \mathrm{~d} t$
with

$$
\begin{equation*}
\theta(t)=\int \frac{\mathrm{d} t}{\eta^{2}} . \tag{32}
\end{equation*}
$$

We see from the scaling of the wavefunctions, equations (17), that the spatial integral in $D_{\mu \nu}^{m n}$ scales as

$$
\begin{equation*}
\int_{-\infty}^{+\infty} u_{\mu} u_{\nu}\left(u_{m} v_{n}\right)^{\prime} \mathrm{d} x \propto \frac{1}{\eta^{2}(t)}=\dot{\theta}(t) . \tag{33}
\end{equation*}
$$

We employ $\theta$ as an integration variable in the temporal integral in (31), which is always possible, because $\dot{\theta}=\eta^{-2}>0$. We see that $D_{\mu \nu}^{m n}$ vanishes unless

$$
\begin{equation*}
\omega_{m}-\omega_{n}=\omega_{\mu}-\omega_{\nu} \tag{34}
\end{equation*}
$$

This frequency constraint is one cornerstone of the orthogonality proof for the $\left(u_{m} v_{n}\right)^{\prime}$ on the products of wavefunctions $u_{\mu} u_{v}$. In addition, we note that the amplitudes $u_{n}$ and $v_{n}$ are solutions of the Schrödinger equation (13) with the effective potential $U+\frac{1}{2} S^{\prime 2}+\dot{S}$ and the eigenfrequencies $\omega_{n}$. Then we apply the reconstruction theorem [1] of one-dimensional Schrödinger equations to prove that

$$
\begin{equation*}
D_{\mu \nu}^{m n} \propto \delta_{\mu m} \delta_{\nu n} \tag{35}
\end{equation*}
$$

We refer the reader to [3, section 3] for an extensive review of the underlying mathematics.
Finally, we turn to the normalization $D_{m n}^{m n}=1$. We perform a three-line calculation starting from equations (20) and (21), and, observing $W_{n}=u_{n}^{\prime} v_{n}-u_{n}^{\prime} v_{n}$ and $W_{n}^{\prime}=0$, obtain

$$
\begin{align*}
D_{m n}^{m n} & =\lim _{T \rightarrow \infty} T^{-1} \int_{-T / 2}^{+T / 2} \int_{-\infty}^{+\infty} u_{m} u_{n}\left(u_{m} v_{n}\right)^{\prime} \mathrm{d} x \mathrm{~d} t \\
& =\lim _{T \rightarrow \infty} T^{-1} \int_{-T / 2}^{+T / 2} \int_{-\infty}^{+\infty}\left[W_{n} u_{n}^{2}+\left(u_{m} u_{n}\right)^{\prime} u_{m} v_{n}\right] \mathrm{d} x \mathrm{~d} t \\
& =\lim _{T \rightarrow \infty} T^{-1} \int_{-T / 2}^{+T / 2} W_{n} \mathrm{~d} t-D_{m n}^{m n} \tag{36}
\end{align*}
$$

We see from (36) that the time average of the Wronskian $W_{n}$ is twice the overlap integral $D_{m n}^{m n}$. The normalization $D_{m n}^{m n}=1$ thus implies that

$$
\begin{equation*}
\lim _{T \rightarrow \infty} T^{-1} \int_{-T / 2}^{+T / 2} W_{n} \mathrm{~d} t=2 \tag{37}
\end{equation*}
$$

This generalizes the Wronskian condition (5) to our case of explicitly time-dependent potentials.

In summary, state reconstruction of one-dimensional wave packets moving in timedependent potentials is possible but restricted to a particular class (28) of potentials, if we require a similar structure (8) as in the time-independent case [1-4]. In essence, the allowed potentials $U(x, t)$ are appropriately scaled and shifted stationary potentials $V(\xi)$ with an additional quadratic term. The time dependence is brought about by an arbitrary scaling $\eta(t)$ and shift $\zeta(t)$. Of course, our result does not prove that otherwise state reconstruction is impossible, but a radically different approach is required. In fact, Johansen [11, 20] has recently found a hydrodynamical method for determining the state of a wave packet that moves in a (potentially) time-dependent potential. Here the density matrix $\left\langle x+y, t_{0}\right| \hat{\rho}\left|x-y, t_{0}\right\rangle$ in position representation is reconstructed from the assumed knowledge of all temporal derivatives of the probability distribution $\operatorname{pr}\left(x, t_{0}\right)$ at a certain time $t_{0}$. This is equivalent to knowledge of the total spatio-temporal evolution $\operatorname{pr}(x, t)$, i.e. to our case, if and only if $\operatorname{pr}(x, t)$ is an analytic function in $t$ on the real axis. Furthermore, Johansen [11] represents the density matrix $\left\langle x+y, t_{0}\right| \hat{\rho}\left|x-y, t_{0}\right\rangle$ as a power series $\sum_{n=0}^{\infty} f_{n}\left(x, t_{0}\right)(n!)^{-1}(2 \mathrm{i} y)^{n}$, assuming $\left\langle x+y, t_{0}\right| \hat{\rho}\left|x-y, t_{0}\right\rangle$ to be analytic in $y$ for all values of $x$.

We may speculate that quantum chaos might limit quantum-mechanical state reconstructions. Why? Classical state measurements are difficult for a chaotic system. The evolution of the system may show little (or a drastic) influence of the initial conditions. Therefore we would expect problems in quantum state determinations due to quantum signatures of chaos [21]. Expressed in terms of the hydrodynamic approach, chaos may render the probability distribution $\operatorname{pr}(x, t)$ and the density matrix $\langle x+y, t| \hat{\rho}|x-y, t\rangle$ less analytic. Note also that our class of allowed potentials correspond to regular motions of wave packets, because the $U(x, t)$ support a complete set of wavefunctions (17) that are scaled
energy eigenfunctions of the potential $V(\xi)$. Therefore, to close this paper with a speculation, the onset of quantum chaos may indeed limit our fundamental ability to infer quantum states from moving wave packets.

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## Appendix A

In this appendix we determine the structure of the phase factors $S_{n}$ that is compatible with the existence of a Wronskian pair (8). We will see that the probability conservation (14) imposes a serious constraint. Let us start from the relation (12) between the irregular and regular wavefunctions

$$
\begin{equation*}
v_{n}=W_{n} u_{n} \int \frac{\mathrm{~d} x}{u_{n}^{2}} \tag{A.1}
\end{equation*}
$$

We take the first and the second spatial derivative, and get

$$
\begin{align*}
& v_{n}^{\prime}=u_{n}^{\prime} W_{n} \int \frac{\mathrm{~d} x}{u_{n}^{2}}+\frac{W_{n}}{u_{n}}  \tag{A.2}\\
& v_{n}^{\prime \prime}=u_{n}^{\prime \prime} W_{n} \int \frac{\mathrm{~d} x}{u_{n}^{2}} . \tag{A.3}
\end{align*}
$$

Therefore

$$
\begin{equation*}
-\frac{1}{2} v_{n}^{\prime \prime}+\left(U+\frac{1}{2} S_{n}^{\prime 2}+\dot{S}_{n}\right) v_{n}=0 \tag{A.4}
\end{equation*}
$$

is satisfied, because $u_{n}$ solves the identical equation (13). Let us see what

$$
\begin{equation*}
\dot{v}_{n}+\frac{1}{2} S_{n}^{\prime \prime} v_{n}+S_{n}^{\prime} v_{n}^{\prime}=0 \tag{A.5}
\end{equation*}
$$

i.e. the probability conservation for $v_{n}$, requires. We differentiate the irregular wavefunction (A.1) with respect to the time $t$, get

$$
\begin{equation*}
\dot{v}_{n}=\dot{u}_{n} W_{n} \int \frac{\mathrm{~d} x}{u_{n}^{2}}+u_{n} \frac{\partial}{\partial t} W_{n} \int \frac{\mathrm{~d} x}{u_{n}^{2}} \tag{A.6}
\end{equation*}
$$

and therefore

$$
\begin{equation*}
\dot{v}_{n}+\frac{1}{2} S_{n}^{\prime \prime} v_{n}+S_{n}^{\prime} v_{n}^{\prime}=u_{n}\left(\frac{\partial}{\partial t} W_{n} \int \frac{\mathrm{~d} x}{u_{n}^{2}}+S_{n}^{\prime} \frac{W_{n}}{u_{n}^{2}}\right) \tag{A.7}
\end{equation*}
$$

The right-hand side of this equation vanishes if

$$
\begin{equation*}
\frac{\partial}{\partial t} W_{n} \int \frac{\mathrm{~d} x}{u_{n}^{2}}+S_{n}^{\prime} \frac{W_{n}}{u_{n}^{2}}=0 \tag{A.8}
\end{equation*}
$$

We take the spatial derivative of this equation and use the probability conservation (14) for the regular wavefunctions to obtain

$$
\begin{align*}
0 & =\frac{\partial}{\partial t}\left(\frac{W_{n}}{u_{n}^{2}}\right)+S_{n}^{\prime \prime} \frac{W_{n}}{u_{n}^{2}}+S_{n}^{\prime}\left(\frac{W_{n}}{u_{n}^{2}}\right)^{\prime} \\
& =\frac{1}{u_{n}^{3}}\left(\dot{W}_{n} u_{n}-2 W_{n} \dot{u}_{n}+S_{n}^{\prime \prime} W_{n} u_{n}-2 S_{n}^{\prime} W_{n} u_{n}^{\prime}\right) \\
& =\frac{1}{u_{n}^{2}}\left(\dot{W}_{n}+2 S_{n}^{\prime \prime} W_{n}\right) \tag{A.9}
\end{align*}
$$

and, consequently,

$$
\begin{equation*}
S_{n}^{\prime \prime}=-\frac{1}{2} \frac{\dot{W}_{n}}{W_{n}} \quad \text { with } \quad W_{n}^{\prime}=0 \tag{A.10}
\end{equation*}
$$

The phases $S_{n}$ are restricted to be quadratic functions in $x$.

## Appendix B

In this appendix we determine the wavefunctions $u_{n}$ that satisfy

$$
\begin{equation*}
\dot{u}_{n}+\frac{1}{2} a u_{n}+(a x+b) u_{n}^{\prime}=0 \tag{B.1}
\end{equation*}
$$

We try the scaling ansatz

$$
\begin{equation*}
u_{n}=\eta^{-1 / 2} w_{n}(\xi) \quad \xi=x / \eta(t)-\zeta(t) \tag{B.2}
\end{equation*}
$$

and differentiate

$$
\begin{align*}
& \dot{u}_{n}=-\frac{1}{2} \frac{\dot{\eta}}{\eta} u_{n}-\eta^{-1 / 2} \frac{\mathrm{~d} w_{n}(\xi)}{\mathrm{d} \xi}\left(\frac{\dot{\eta}}{\eta^{2}} x+\dot{\zeta}\right)  \tag{B.3}\\
& u_{n}^{\prime}=\eta^{-1 / 2} \frac{\mathrm{~d} w_{n}(\xi)}{\mathrm{d} \xi} \frac{1}{\eta} \tag{B.4}
\end{align*}
$$

The term in (B.1) that is proportional to $u_{n}$ vanishes if

$$
\begin{equation*}
a=\frac{\dot{\eta}}{\eta} \tag{B.5}
\end{equation*}
$$

The term that contains the derivative $\mathrm{d} w_{n} / \mathrm{d} \xi$ is zero if, in addition,

$$
\begin{equation*}
b=\eta \dot{\zeta} \tag{B.6}
\end{equation*}
$$

In this way we have found the general solution of (B.1).

## Appendix C

In this appendix we determine the scaling of

$$
\begin{equation*}
F_{n}=U+\frac{1}{2} S^{\prime 2}+\dot{S}-\dot{\Omega}_{n} \tag{C.1}
\end{equation*}
$$

In terms of the $F_{n}$ the Schrödinger equation (13) reads as

$$
\begin{equation*}
u_{n}^{\prime \prime}=2 F_{n} u_{n} \tag{C.2}
\end{equation*}
$$

Therefore

$$
\begin{equation*}
\dot{u}_{n}^{\prime \prime}=2\left(\dot{F}_{n} u_{n}+F_{n} \dot{u}_{n}\right)=2\left(\dot{F}_{n} u_{n}-\frac{1}{2} S^{\prime \prime} F_{n} u_{n}-S^{\prime} F_{n} u_{n}^{\prime}\right) \tag{C.3}
\end{equation*}
$$

using the probability conservation (14). On the other hand, we take the second spatial derivative of (14), observe $S^{\prime \prime \prime}=0$, and get

$$
\begin{equation*}
\dot{u}_{n}^{\prime \prime}=-\frac{5}{2} S^{\prime \prime} u_{n}^{\prime \prime}-S^{\prime} u_{n}^{\prime \prime \prime}=-5 S^{\prime \prime} F_{n}-2 S^{\prime} F_{n} u_{n}^{\prime}-2 S^{\prime} F_{n}^{\prime} u_{n} \tag{C.4}
\end{equation*}
$$

We combine equations (C.3) and (C.4) to obtain

$$
\begin{equation*}
\dot{F}_{n}+2 S^{\prime \prime} F_{n}+S^{\prime} F_{n}^{\prime}=0 \tag{C.5}
\end{equation*}
$$

or, via the quadratic structure (16) of the phase $S$,

$$
\begin{equation*}
\dot{F}_{n}+2 a F_{n}+(a x+b) F_{n}^{\prime}=0 \tag{C.6}
\end{equation*}
$$

We proceed along similar lines as in appendix B and finally obtain

$$
\begin{equation*}
F_{n}=\eta^{-2} G_{n}(\xi) \quad \xi=x / \eta(t)-\zeta(t) \tag{C.7}
\end{equation*}
$$

In this way we determined the scaling of the right-hand side of the Schrödinger equation (13), if we impose the existence of Wronskian pairs (8).

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    $\ddagger$ After submission of this paper, a letter [11] appeared that describes a hydrodynamical method for state reconstruction of wave packets that travel in potentially time-dependent potentials.

